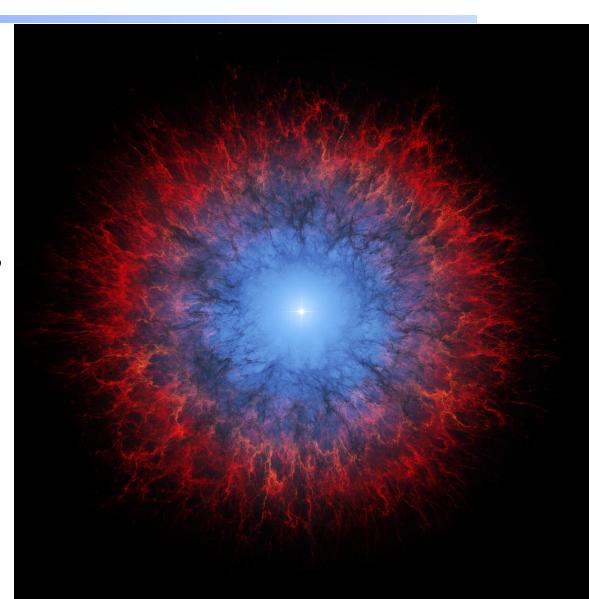
Velocity fields

- Default is static, with thermal broadening
- Turbulence can be added
 - makes line optical depths smaller, so lines escape more easily, continuum florescent excitation more important

- Winds can also be computed
- Line transfer with "Large Velocity Gradient" (LVG) or "Sobolev approximation"
 - -2 names for same thing

Wind solutions

- Cloudy will solve for the wind structure
- Wind ballistic supersonic outflows
 - Positive wind velocity
- dynamics_wind.in



Wind solutions

- ~sonic flows from H II regions
 - Negative velocity, since motion is towards star
- ◆ D-critical flows, nearly at speed of sound
- dynamics orion flow.in
- Described <u>here</u> and <u>here</u>

Project poster

- One page landscape format PDF with results of the project
- One per group, to be posted on the web site

- **◆** Title, authors, abstract
- Introduction
 - What problem were you trying to solve?
- Methods and calculations
- Conclusions
- Due by July 1

Some closing thoughts

• Quantitative spectroscopy - read the message in the starlight – what does the spectrum tells us?

- ◆ Like all fields, a steep learning curve, but the rewards will be great - be able to decipher the message
 - Like medieval priests, an elevated position since only a few can read the sacred texts