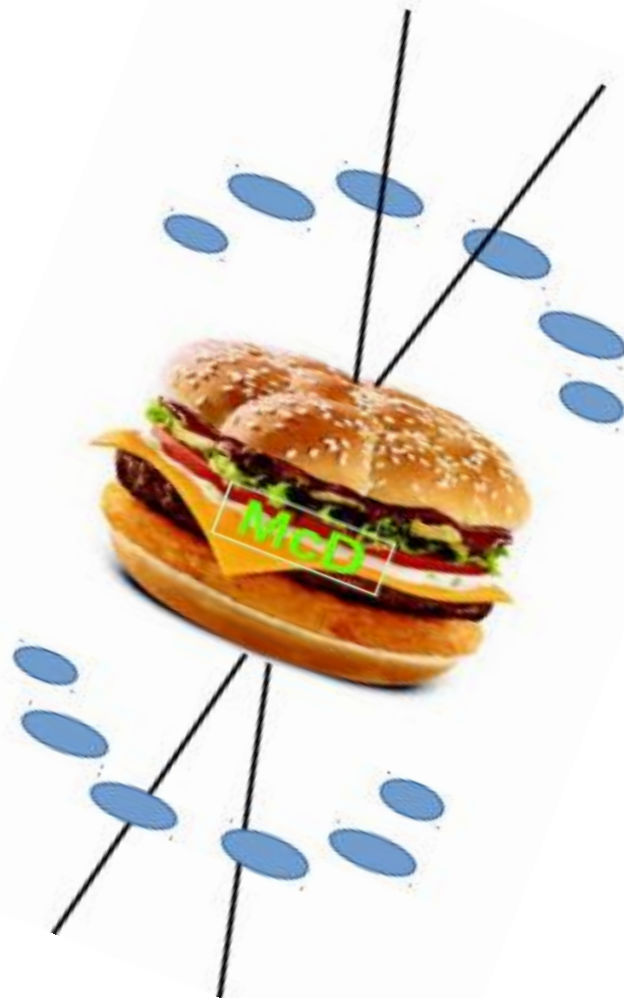


Modelling of C IV absorption lines in BAL QSOs using CLOUDY McD

Sapna , Savithri, Aditi, Navpreet,
Parveen, Ritesh, Vineet

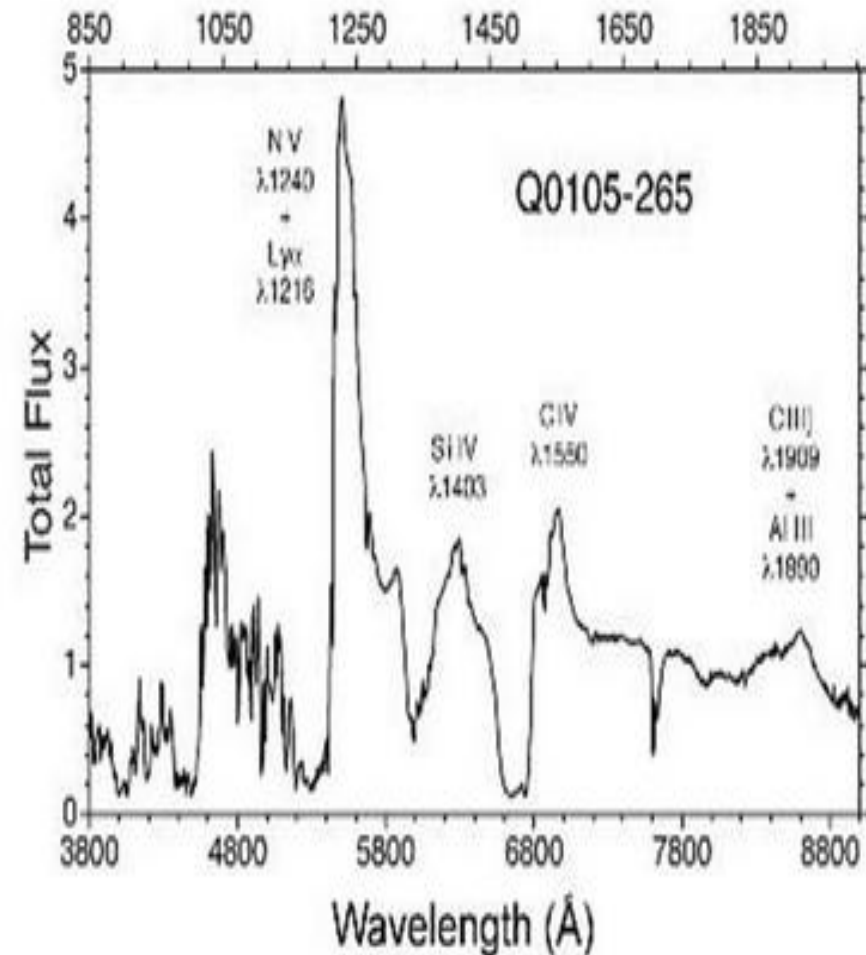
Cloudy Workshop
(21-26) Sept., 2015
(IUCAA, Pune)

Special Thanks:
Susmita Chakravorty
Hum Chand
Gary J. Ferland



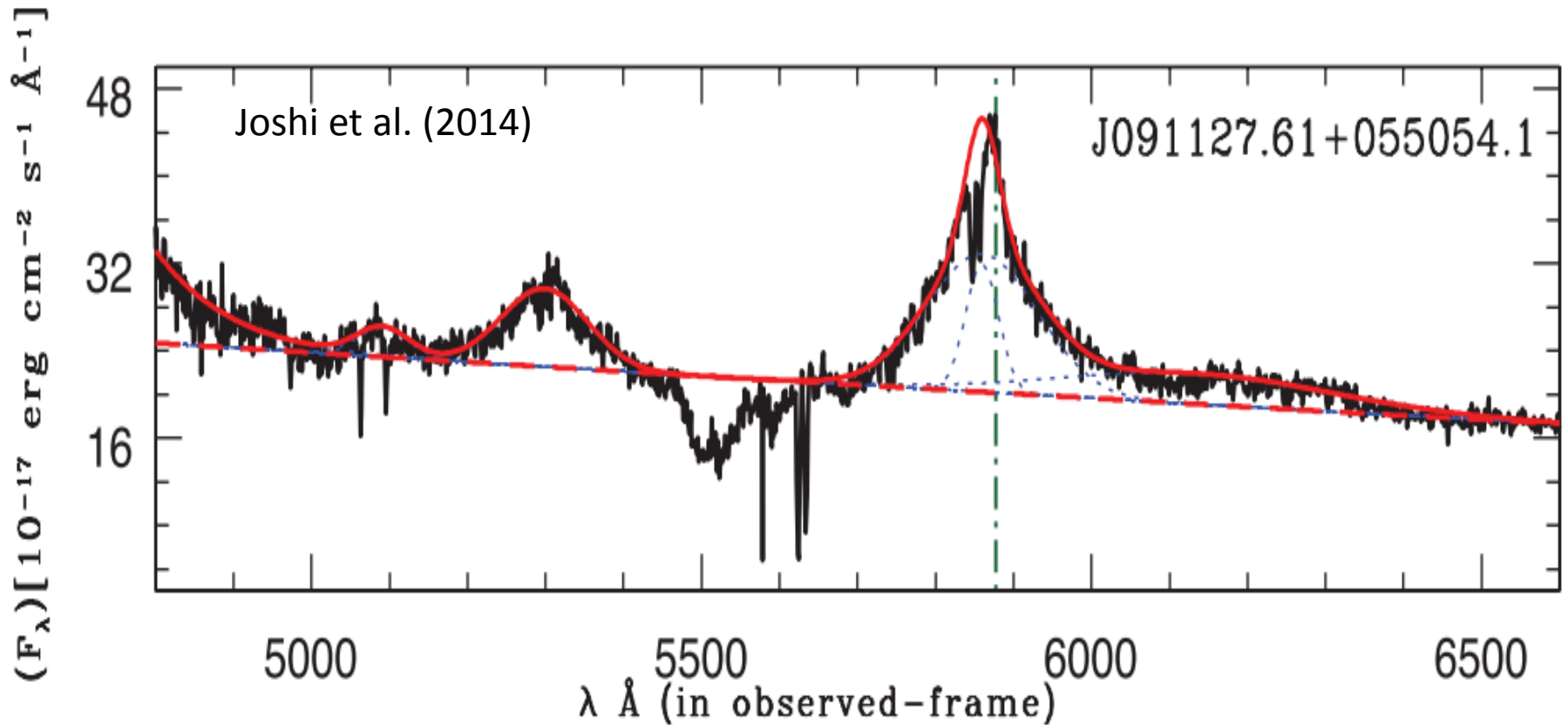
BAL QSO : Broad Absorption Line QSOs

- BAL QSOs comprise about 3-10% of QSO population
- Shows strong blue-shifted broad absorption lines
- Strength of the BAL troughs change with time
- Outflows are signature of AD
- Outflow velocities $10\text{-}30\text{k km s}^{-1}$
- Studies of BAL trough intensity and outflow velocity variability are rare



J091127+055054

- Gravitationally lensed QSO
- X-ray luminosity $\sim 4 \times 10^{46}$ ergs s^{-1}
- $Z = 2.793$, $\alpha(\text{ox}) = -1.58$

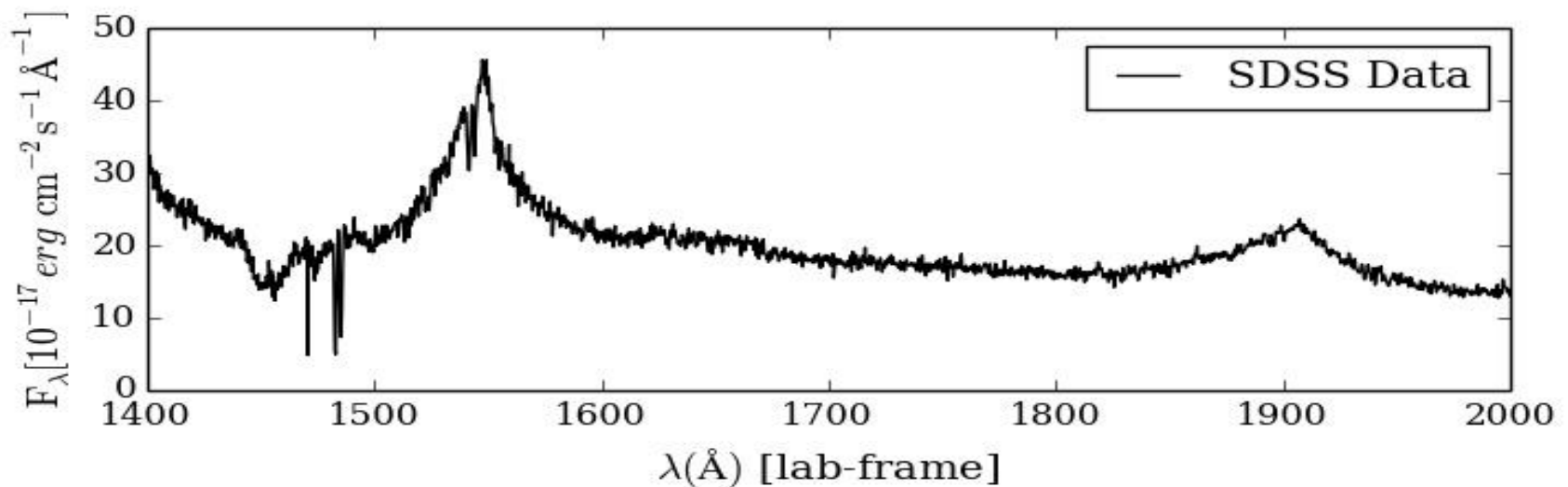
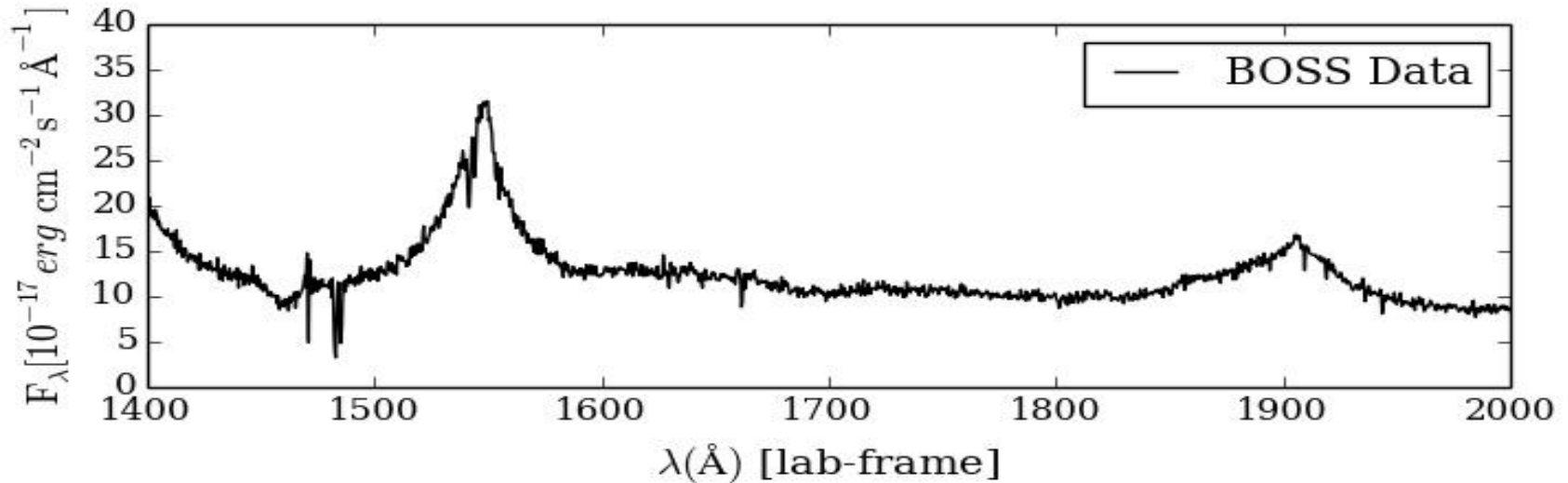


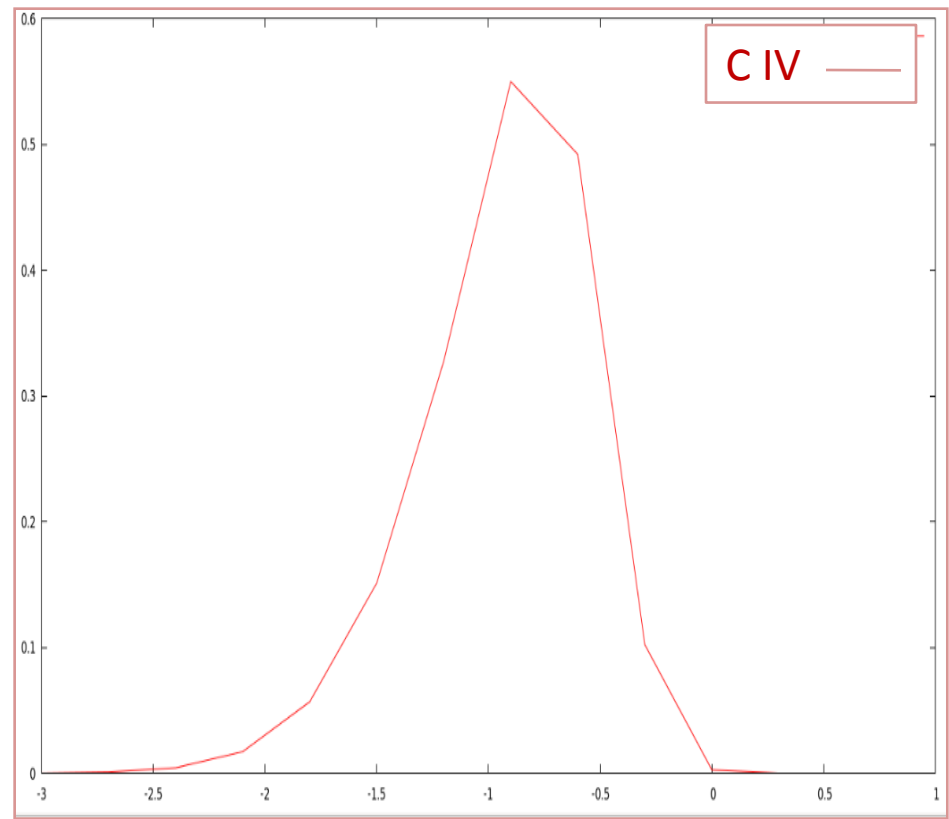
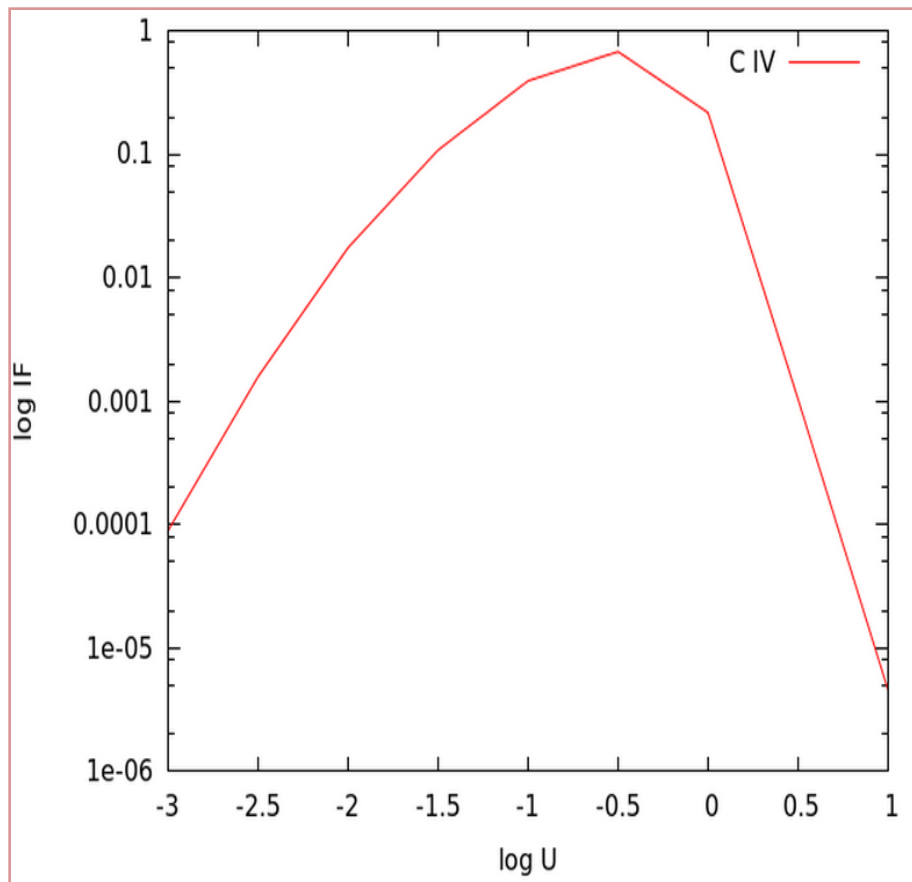
Motivation

- Observed average deceleration for whole BAL profile $\sim -2.0 \pm 0.1 \text{ cm s}^{-2}$ over time-span of 2-3 years.
- Physical conditions, acceleration mechanisms, location of QSO outflows are poorly understood.
- Line variability helps probing the structure and dynamics of these out flowing gas.

“Cloudy” as a tool!

Spectra obtained from BOSS & SDSS data

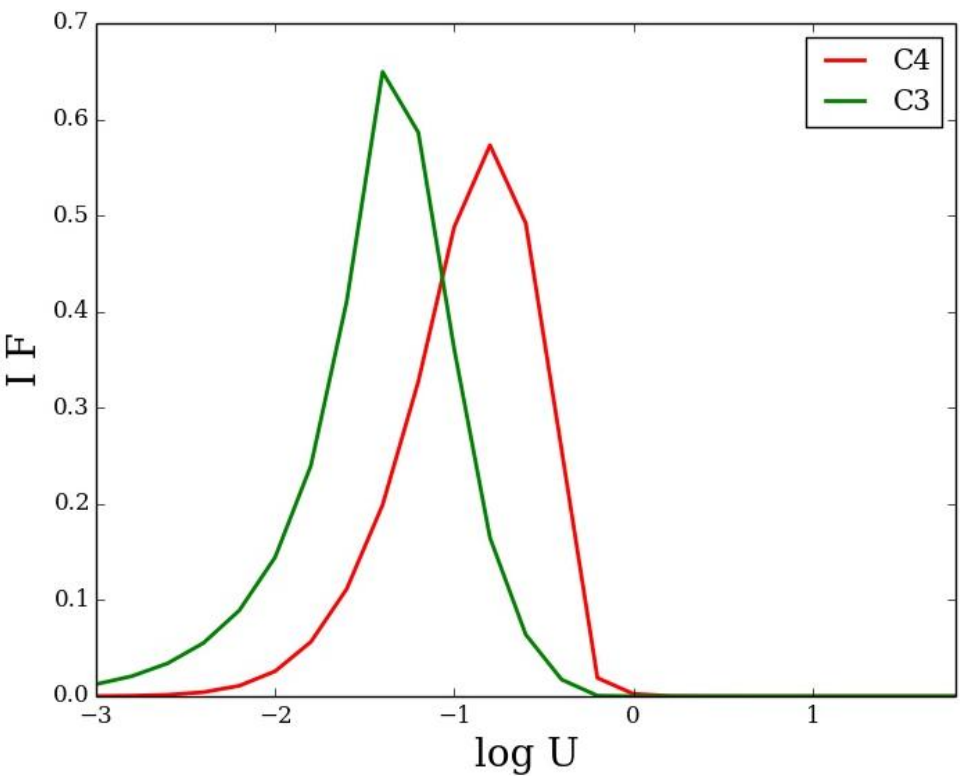




Plot between $\log (U)$ and $\log (I.F.)$

Value at which I.F. is maximum will be given as input for SED modelling.

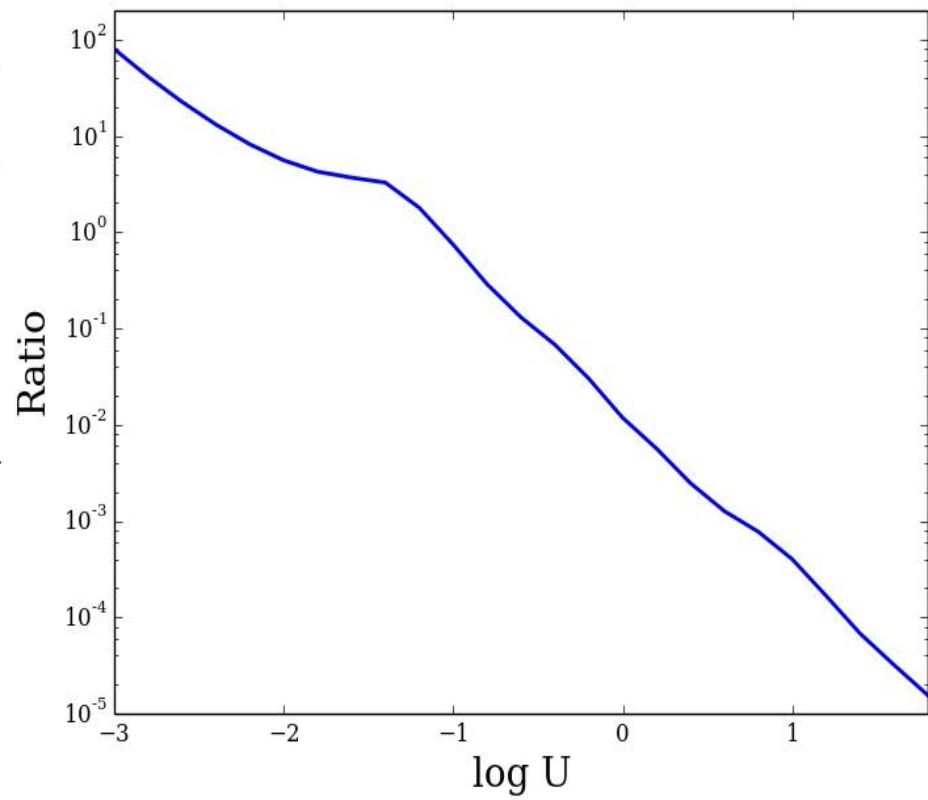
Model SED for which CIII/CIV ratio best matches with the observed SED



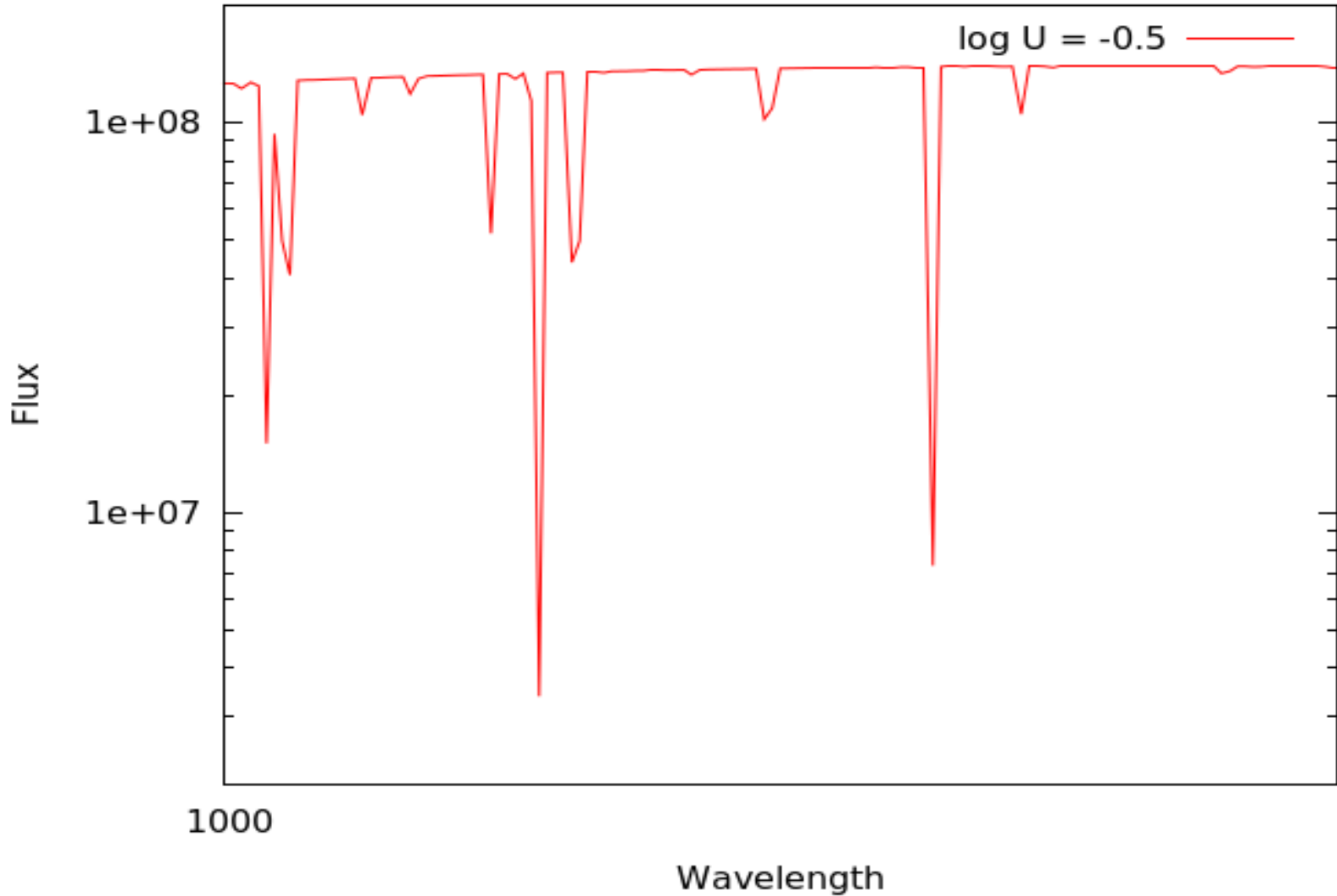
Comparison of I.F.
For C3 and C4 lines.



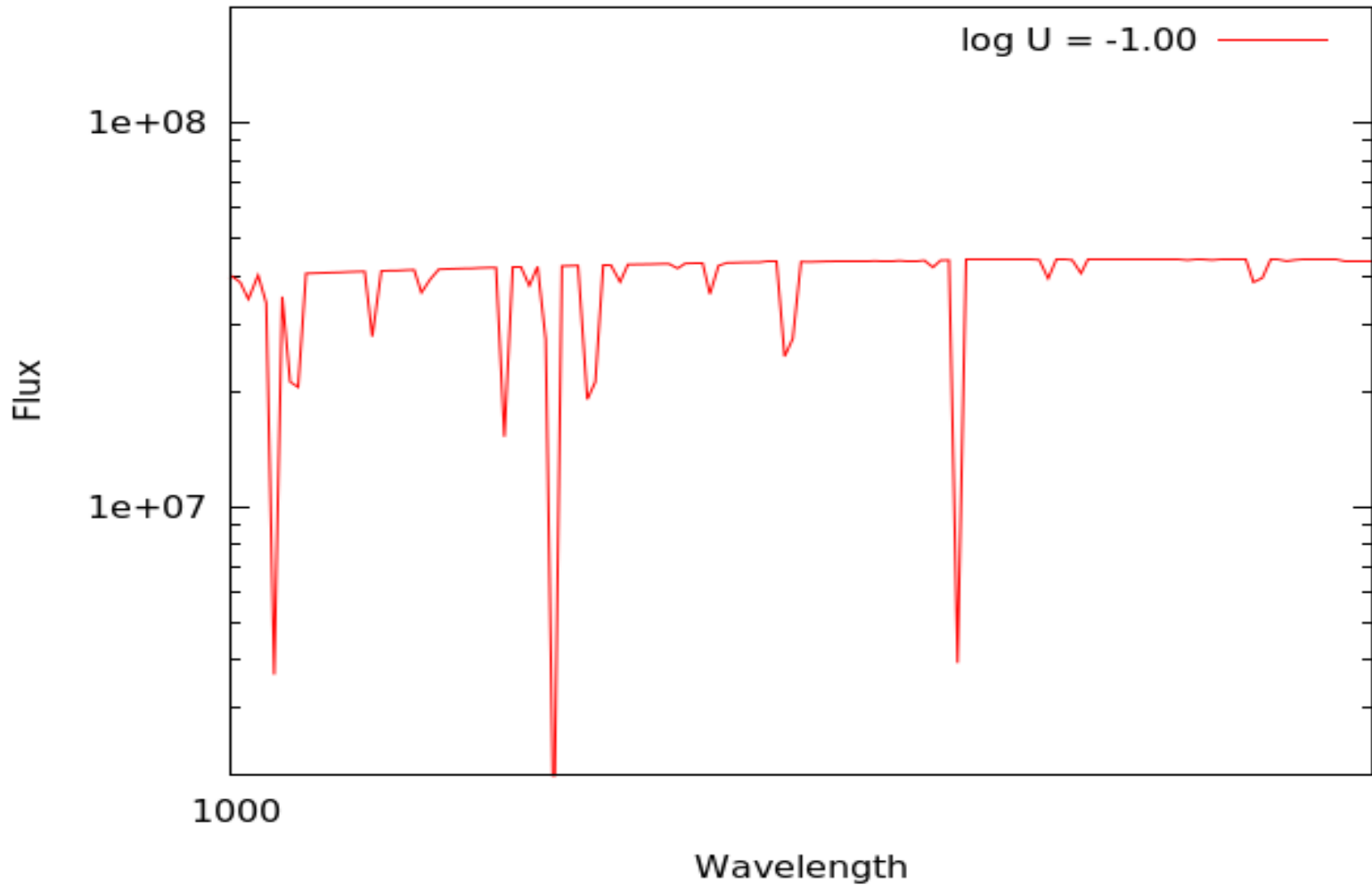
Ratio of C3 to C4



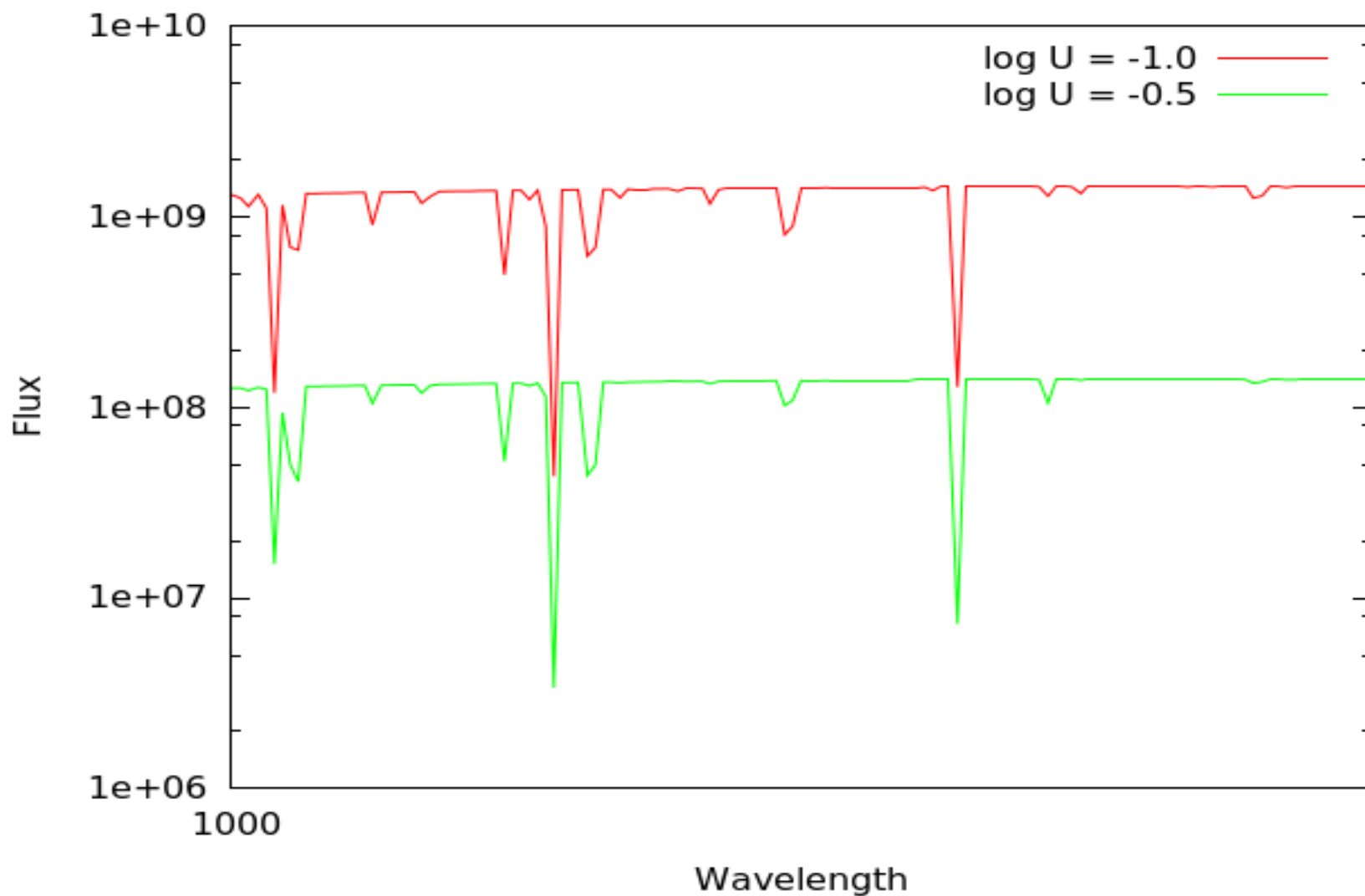
Modelled spectra for $\log U = -0.5$



Modelled spectra for $\log U = -1.0$



Over plotted the spectra for $\log U = -0.5$, -1.0



$$U = \frac{Q(H)}{4\pi n_H r^2 c}$$

U = Ionization parameter

R = separation [cm] between the center of the source and illuminated face of the cloud.

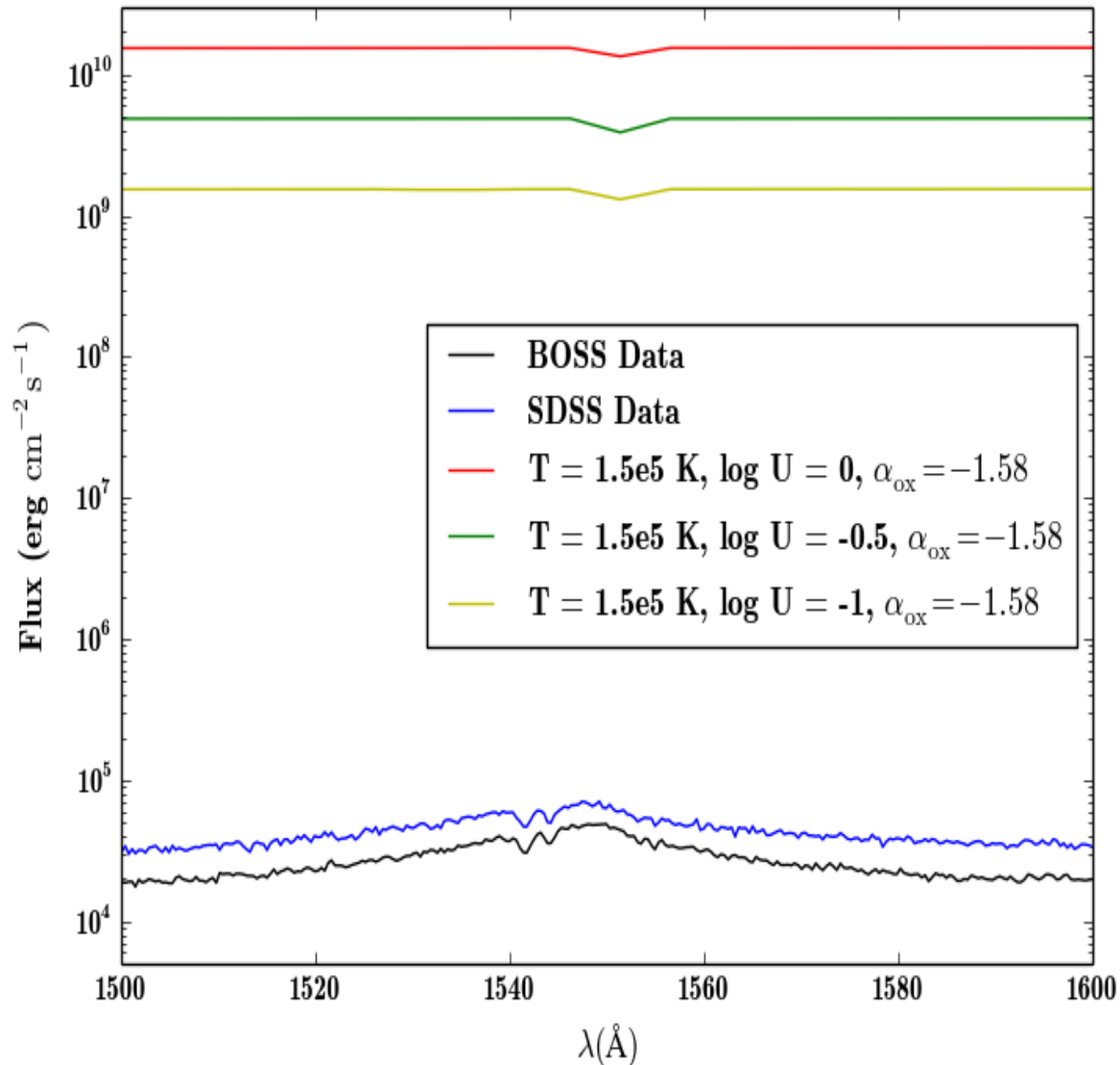
n(H) = total hydrogen density

c = the speed of light

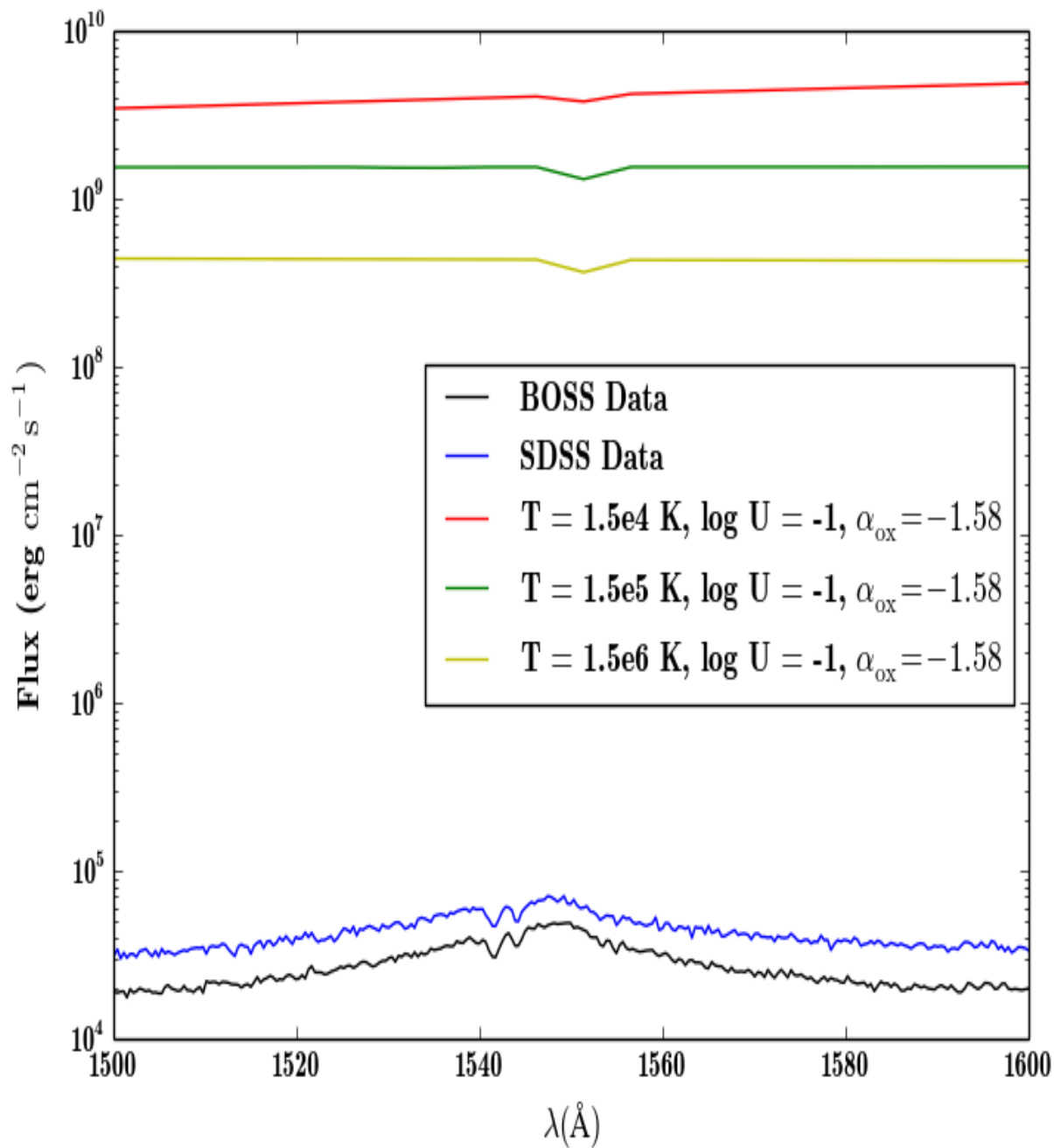
Q(H) = number of hydrogen-ionizing photons

For two different U values which we calculated for the graph, variation in Q(H) can be deduced

RESULTS

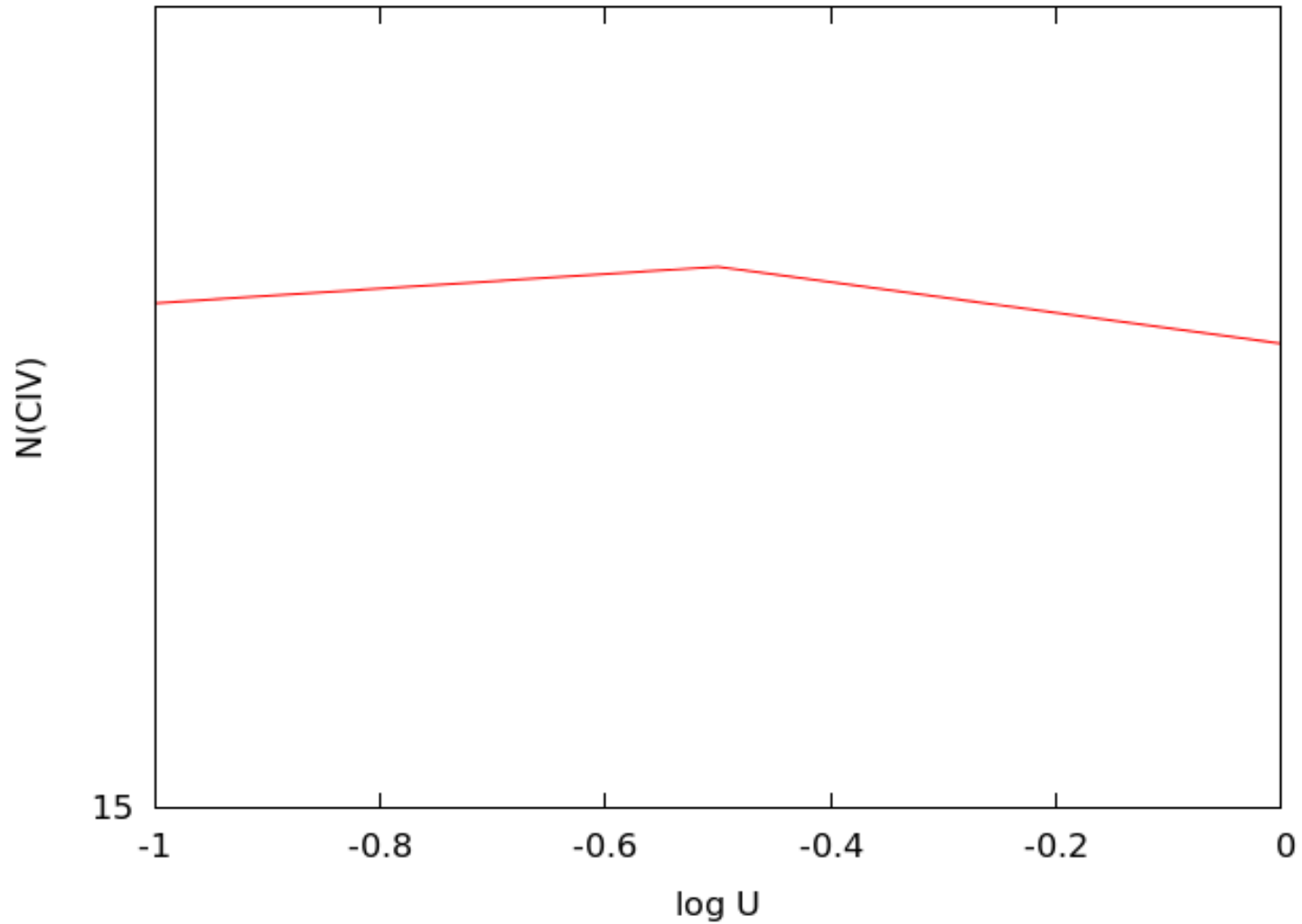


Log U	N (CIV) (in log scale)
0	17.25 cm ⁻²
-0.5	18.22 cm ⁻²
-1	17.78 cm ⁻²

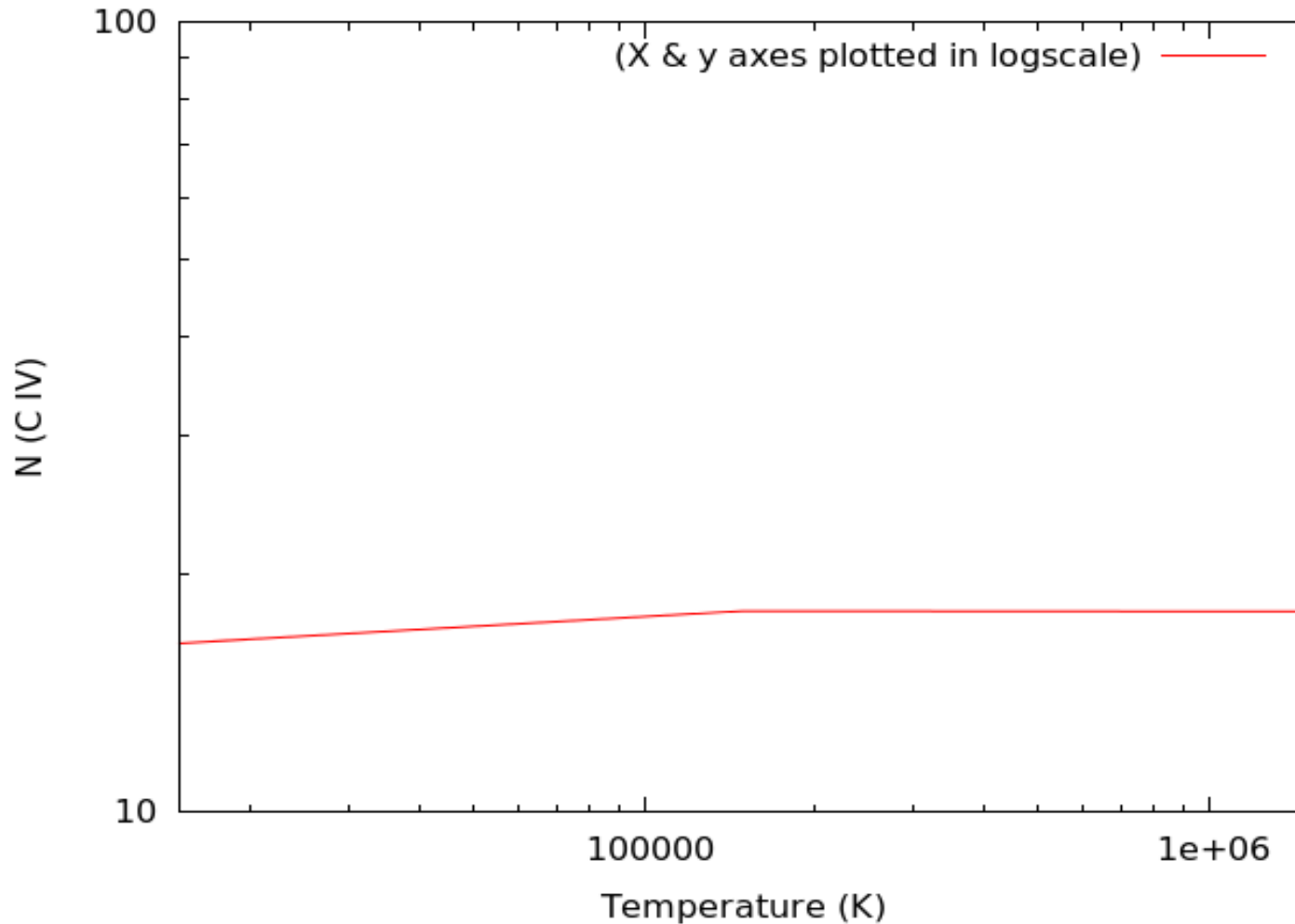


T (in K)	N (CIV) (in log scale)
1.5×10^4	16.367 cm^{-2}
1.5×10^5	17.983 cm^{-2}
1.5×10^6	17.965 cm^{-2}

*Plot showing variation of
N (CIV) with Ionization parameter*



*Plot showing variation of
N (CIV) with Ionizing photon Temperature*



Observed Results

$$N1 \text{ (at 55896 MJD)} = 2.92374 \times 10^{14} \text{ cm}^{-2}$$

$$N2 \text{ (at 52652 MJD)} = 3.8307182 \times 10^{14} \text{ cm}^{-2}$$

$$(N1 - N2) \sim 9.0692 \times 10^{13} \text{ cm}^{-2}$$

$$N1/N2 \sim 1.3102078$$

Calculated Results

$$N1 (\log U = -1)/N2 (\log U = 0) = 1.8113 \text{ (at same T)}$$

WHAT COULD BE POSSIBLY DONE..??

To match modelled data with the observed data
with possible accuracy

THANKS